

Lecture 4

The Principle of Equivalence

What is gravity?

Uniformly Accelerated Motion

World-line $w(\lambda)$

$$\Rightarrow \text{4-velocity } v(\lambda) = \frac{dw}{d\tau}(\lambda)$$

$$\Rightarrow \text{4-acceleration } a(\lambda) = \frac{dv}{d\tau}(\lambda)$$

$$\Rightarrow \text{4-jerk } j(\lambda) = \frac{da}{d\tau}(\lambda)$$

Newtonian physics: $j=0$ \leftarrow Does not work in SR!

$$v(\lambda) \cdot v(\lambda) = -c^2$$

$$\Rightarrow v(\lambda) \cdot a(\lambda) = 0$$

$$\Rightarrow a(\lambda) \cdot a(\lambda) + v(\lambda) \cdot j(\lambda) = 0$$

$a(\lambda)$ is space-like, so unless $a(\lambda) = 0$, $\|a(\lambda)\|^2 > 0$

$\Rightarrow j(\lambda) \neq 0 \leftarrow$ must have a component along $v(\lambda)$

Uniform acceleration in SR
 makes the minimal assumption
 that $j(\lambda) \propto v(\lambda)$

$$j(\lambda) = \alpha v(\lambda)$$

$$\Rightarrow \|a(\lambda)\|^2 + \alpha \|v(\lambda)\|^2 = 0$$

$$\Rightarrow \alpha = \frac{\|a(\lambda)\|^2}{c^2}$$

$$\text{So } j(\lambda) = \frac{\|a(\lambda)\|^2}{c^2} v(\lambda)$$

$$\Rightarrow a(\lambda) \cdot j(\lambda) = 0$$

$$\parallel$$

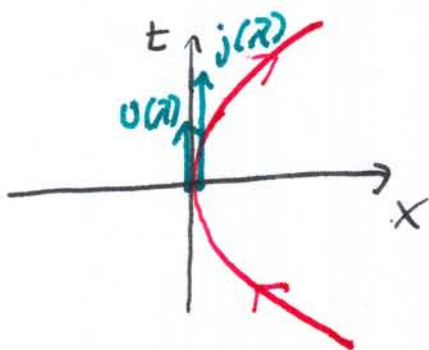
$$\frac{1}{2} \frac{d}{d\tau} \|a(\lambda)\|^2$$

$$\Rightarrow \|a(\lambda)\|^2 = g^2 = \text{const.}$$

↑ positive b/c
 $a(\lambda)$ space-like

$$\Rightarrow j(\lambda) = \frac{d^2}{d\tau^2} v(\lambda) = \frac{g^2}{c^2} v(\lambda)$$

Note: In uniform acceleration, the spatial projection of $\vec{j}(\lambda)$ vanishes in the instantaneously co-moving frame: $\vec{j}(\lambda) = 0$



$$\vec{j}(\lambda) = \frac{d^3}{d\tau^3} \vec{x}(\lambda)$$

$$= \left(\frac{dt}{d\tau} \frac{d}{dt} \right)^3 \vec{x}(\lambda)$$

$$\uparrow \quad \uparrow$$

$$\gamma(\lambda) = \frac{1}{\sqrt{1 - v^2(\lambda)/c^2}}$$

$$\left(\gamma \frac{d}{dt} \right)^3 = \gamma^3 \frac{d^3}{dt^3} + 3 \gamma^2 \dot{\gamma} \frac{d^2}{dt^2}$$

$$+ \gamma (\dot{\gamma}^2 + \gamma \ddot{\gamma}) \frac{d}{dt} \quad \begin{array}{l} = 0 \text{ in} \\ \text{icio} \end{array}$$

$$\vec{j}(\lambda) = \gamma^3(\lambda) \frac{d^3 \vec{x}}{dt^3}(\lambda) + 3 \gamma^2(\lambda) \dot{\gamma}(\lambda) \frac{d^2 \vec{x}}{dt^2}(\lambda)$$

$$+ \gamma(\lambda) (\dot{\gamma}^2(\lambda) + \gamma(\lambda) \ddot{\gamma}(\lambda)) \frac{d \vec{x}}{dt}(\lambda)$$

$= 0$ in icio \uparrow

\therefore Uniform acceleration



icio sees no (Newtonian) jerk.

Need to solve ODE:

$$\frac{d^2}{d\tau^2} U(\tau) = \frac{g^2}{c^2} U(\tau)$$

$$\begin{aligned} \mapsto U(\tau) = & U_0 \cosh\left(\frac{g}{c}(\tau - \tau_0)\right) \\ & + \frac{c}{g} a_0 \sinh\left(\frac{g}{c}(\tau - \tau_0)\right) \end{aligned}$$

$U_0 =$ 4-velocity at $\tau = \tau_0$

$a_0 =$ 4-acceleration at $\tau = \tau_0$

$$\begin{aligned} \|U(\tau)\|^2 &= \|U_0\|^2 \cosh^2\left(\frac{g}{c}(\tau - \tau_0)\right) \\ &+ \|a_0\|^2 \frac{c^2}{g^2} \sinh^2\left(\frac{g}{c}(\tau - \tau_0)\right) \\ &+ 2 U_0 \cdot a_0 \frac{c}{g} \cosh\left(\frac{g}{c}(\tau - \tau_0)\right) \sinh(\dots) \\ &= -c^2 \cosh^2(\dots) + c^2 \sinh^2(\dots) \\ &= -c^2 \end{aligned}$$

$\Rightarrow U(\tau)$ is proper-time parameterized

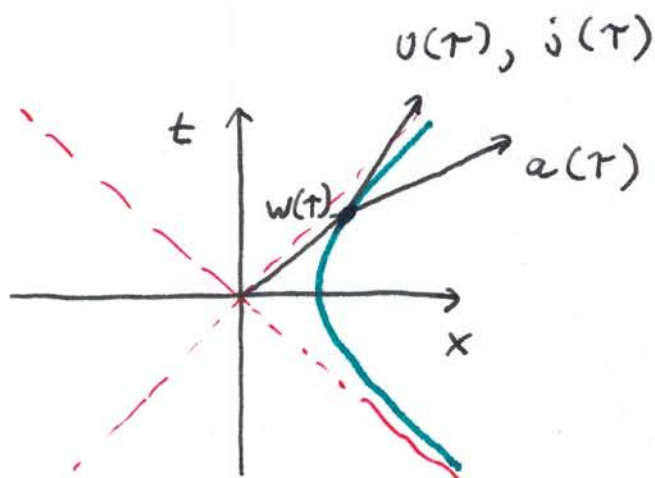
Integrate one more time

$$w(\tau) = w_0 + \frac{c}{g} v_0 \sinh\left(\frac{g}{c}(\tau - \tau_0)\right) + \frac{c^2}{g^2} a_0 \cosh\left(\frac{g}{c}(\tau - \tau_0)\right)$$

Take $w_0 = 0, \tau_0 = 0$

$$\|w(\tau)\|^2 = \frac{c^2}{g^2} \cdot -c^2 \sinh^2(\dots) + \frac{c^4}{g^4} \cdot g^2 \cosh^2(\dots) = \frac{c^4}{g^2}$$

constant

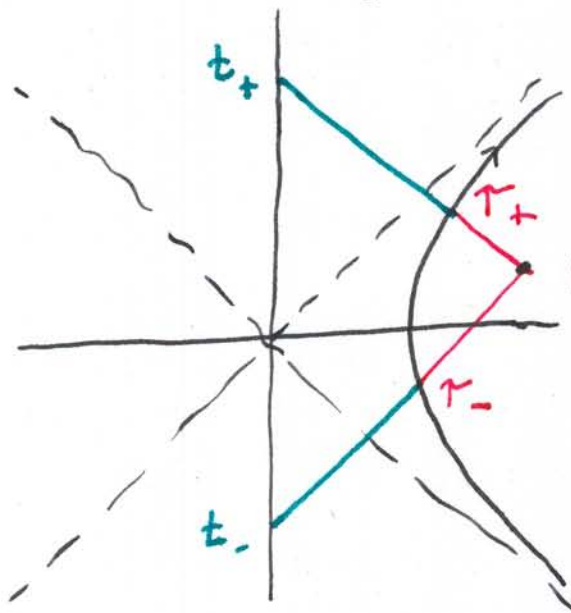


Uniform acceleration is hyperbolic motion in the U - a plane (2-d.)

$$\begin{pmatrix} t \\ x \end{pmatrix} = \begin{pmatrix} \frac{c}{g} \sinh\left(\frac{g}{c}(\tau - \tau_0)\right) \\ \frac{c^2}{g} \cosh\left(\frac{g}{c}(\tau - \tau_0)\right) \end{pmatrix}$$

$$x(t) = \frac{c^2}{g} \sqrt{1 + \left(\frac{g}{c}t\right)^2}$$

Radio-Coordinates of an Accelerating Observer



$$t = \frac{t_+ + t_-}{2}$$

$$x = \frac{t_+ - t_-}{2} c$$

$$E = (t, x) = (\tau, \xi)$$

$$\tau = \frac{\tau_+ + \tau_-}{2}$$

$$\xi = \frac{\tau_+ - \tau_-}{2} c$$

$$t_+ = t(\tau_+) + \frac{x(\tau_+)}{c}$$

$$= \frac{c}{g} \left[\sinh \frac{g\tau_+}{c} + \cosh \frac{g\tau_+}{c} \right]$$

$$= \frac{c}{g} e^{g\tau_+/c}$$

$$t_- = t(\tau_-) - \frac{x(\tau_-)}{c} = - \frac{c}{g} e^{-g\tau_-/c}$$

$$t = \frac{c}{2g} \left(e^{g\tau_+/c} - e^{-g\tau_-/c} \right) = \frac{c}{g} e^{\frac{g\xi}{c^2}} \sinh \frac{g\tau}{c}$$

$$x = \frac{c^2}{2g} \left(e^{g\tau_+/c} + e^{-g\tau_-/c} \right) = \frac{c^2}{g} e^{\frac{g\xi}{c^2}} \cosh \frac{g\tau}{c}$$

$$\tau_{\pm} = \tau \pm \frac{\xi}{c}$$

Inverse transformation:

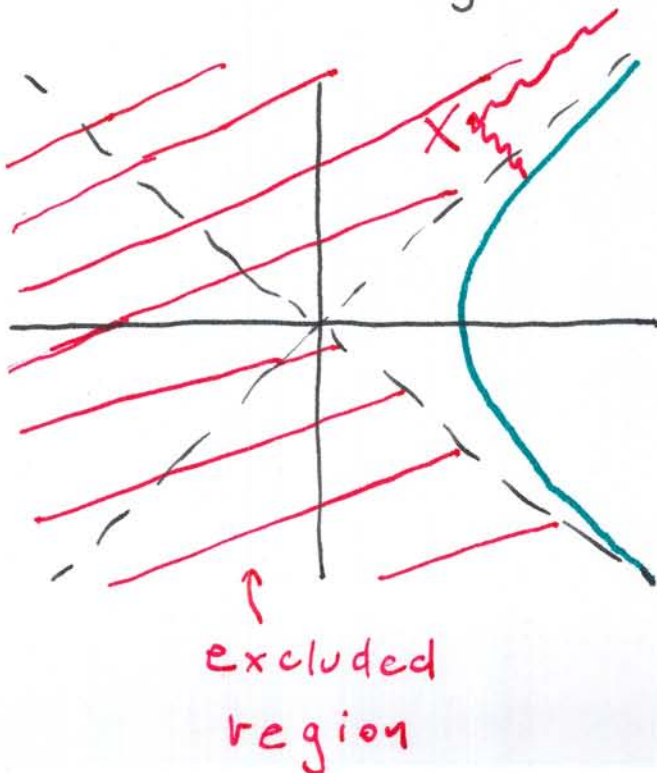
$$t_{\pm} = \pm \frac{c}{g} e^{\pm g \tau_{\pm} / c}$$

$$\Rightarrow \tau_{\pm} = \pm \frac{c}{g} \ln\left(\pm \frac{g}{c} t_{\pm}\right)$$

$$\begin{aligned} \Rightarrow \tau \pm \frac{z}{c} &= \pm \frac{c}{g} \ln\left(\pm \frac{g}{c} \left(t \pm \frac{x}{c}\right)\right) \\ &= \pm \frac{c}{g} \ln\left(\frac{g}{c^2} (x \pm ct)\right) \end{aligned}$$

$$\Rightarrow \tau = \frac{c}{2g} \ln \frac{x+ct}{x-ct}$$

$$z = \frac{c^2}{2g} \ln \left(\frac{g^2}{c^4} (x^2 - c^2 t^2) \right)$$



Radio coordinates cover only the Rindler wedge in Minkowski spacetime.

The metric in accelerating
radio coordinates:

$$t = \frac{c}{g} e^{g\bar{z}/c^2} \sinh\left(\frac{g\bar{T}}{c}\right)$$

$$x = \frac{c^2}{g} e^{g\bar{z}/c^2} \cosh\left(\frac{g\bar{T}}{c}\right)$$

$$-c^2 dt^2 + dx^2 = e^{2g\bar{z}/c^2} (-d\bar{T}^2 + d\bar{z}^2)$$

↑
exercise!